

March 5, 1999

PHYSICAL APPLICATIONS OF GEOMETRIC ALGEBRA

LECTURE 15

SUMMARY

To find the field equations we must establish **covariant** forms of the **field strength tensor**. Dimensional analysis then tells what the correct equations are. Similar considerations apply to **trajectories** and enable us to identify the **metric**.

- The **covariant field strength tensors**.
- The **Planck scale** and the magnitude of the field strength of the displacement field.
- Further properties of the Riemann tensor.
- The second field equation.
- **Trajectories**.
- The **metric** and the link with **GR**.

COVARIANT FIELD STRENGTHS

Need **covariant** forms of field strength. Start with **rotation gauge**. $\Omega(a)$ removes terms in $a \cdot \nabla R \tilde{R}$. Under displacements must have

$$\Omega(a; x) \mapsto \Omega'(a; x) = \Omega[f(a); x']$$

Field strength has term in $\Omega(a) \times \Omega(b)$. Must transform to

$$R(a \wedge b) \mapsto R'(a \wedge b) = R[f(a) \wedge f(b)] = R[f(a \wedge b); x']$$

Picks up a term in $f(B)$. Remove this with suitable version of $\bar{h}(a)$. Has

$$\bar{h}(a) \mapsto \bar{h}'(a) = \bar{h}f^{-1}(a)$$

so adjoint transforms as

$$h(a) \mapsto h'(a) = f^{-1}h(a).$$

Insert this into $R(B)$. Define covariant field strength

$$\mathcal{R}(B) = R[h(B)]$$

Factor of $h(B)$ alters rotation gauge properties.

$$\bar{h}(a) \mapsto \bar{h}'(a) = R\bar{h}(a)\tilde{R}.$$

so adjoint goes as

$$h(a) \mapsto h'(a) = \partial_b \langle a R \bar{h}(b) \tilde{R} \rangle = h(\tilde{R}aR).$$

Summarise transformation properties of $\mathcal{R}(B)$ by:

$$\text{Displacements: } \mathcal{R}'(B, x) = \mathcal{R}(B, x')$$

$$\text{Rotations: } \mathcal{R}'(B) = R\mathcal{R}(\tilde{R}BR)\tilde{R}.$$

Just what we want for a **covariant tensor**. Call $\mathcal{R}(B)$ the **Riemann tensor**. Understand rotation transformation from

$$\mathcal{R}(B) = \alpha B$$

This is 'dilate all fields by factor α '. Transformed field is

$$\mathcal{R}'(B) = R\mathcal{R}(\tilde{R}BR)\tilde{R} = R(\alpha\tilde{R}BR)\tilde{R} = \alpha B$$

Same **physical information**.

DISPLACEMENT GAUGE

Key quantity is

$$\mathcal{S}(a) = -\bar{h}[\nabla \wedge \bar{h}^{-1}(a)] = \bar{h}(\dot{\nabla}) \wedge \dot{\bar{h}}\bar{h}^{-1}(a).$$

But $\bar{h}(a)$ picks up additional rotors under rotation gauge.

Replace the directional derivatives by covariant derivatives:

$$\mathcal{S}(a) = \bar{h}(\partial_b) \wedge \left(b \cdot \dot{\nabla} \dot{\bar{h}}\bar{h}^{-1}(a) + \Omega(b) \cdot a \right).$$

Guarantees required transformation laws

$$\text{Displacements: } \mathcal{S}'(a, x) = \mathcal{S}(a, x')$$

$$\text{Rotations: } \mathcal{S}'(a) = R\mathcal{S}(\tilde{R}aR)\tilde{R}.$$

THE FIELD EQUATIONS

Can get field equations from a **Lagrangian** and action principle. Will use dimensional analysis instead. Have 3 constants G , \hbar and c . Fix the natural scale for interactions:

Planck length $l_P = \sqrt{\hbar G/c^3} = 1.6 \times 10^{-35} \text{ m}$

Planck mass $M_P = \hbar/L_P c = 2.2 \times 10^{-8} \text{ kg}$

Planck time $t_P = L_P/c = 5.3 \times 10^{-44} \text{ s.}$

$\Omega(a)$ has dimensions of $(\text{length})^{-1}$ and $\bar{h}(a)$ is dimensionless. So $\mathcal{S}(a)$ and $\mathcal{R}(B)$ differ in dimensions by a factor of **length**. Expect $\mathcal{S}(a)$ comparable in magnitude to $l_P \mathcal{R}(B)$. **Extremely small!** If ignore quantum effects, expect $\mathcal{S}(a)$ vanishes. Gives first field equation

$$\mathcal{S}(a) = 0.$$

(In fact, from Lagrangian analysis $\mathcal{S}(a)$ driven by **quantum spin**.) Does **not** mean $\bar{h}(a)$ is a pure gauge, because $\Omega(a)$ field couples into covariant field strength! This coupling generates some dynamics. Also ensures that equations are (locally) those of GR!

CONSEQUENCES

First write $\mathcal{S}(a) = 0$ in form

$$\mathcal{D} \wedge \bar{h}(a) = \bar{h}(\partial_b) \wedge [b \cdot \nabla \bar{h}(a) + \Omega(b) \cdot \bar{h}(a)] = 0$$

Version for vector field useful

$$\mathcal{D} \wedge \bar{h}(A) = \bar{h}(\nabla \wedge A)$$

Removes a potential ambiguity. Electromagnetic vector potential is A . Covariant version is

$$\mathcal{A} = \bar{h}(A)$$

(Because A generalises $\nabla \phi$). Could generate electromagnetic field strength from A or \mathcal{A} . But now both give the **same** result

$$\mathcal{F} = \bar{h}(\nabla \wedge A) = \mathcal{D} \wedge \mathcal{A} = \bar{h}(F).$$

Now extend to a bivector,

$$\begin{aligned} \mathcal{D} \wedge \bar{h}(a \wedge b) &= \bar{h}(\partial_c) \wedge [\mathcal{D}_c \bar{h}(a) \wedge \bar{h}(b) + \bar{h}(a) \wedge \mathcal{D}_c \bar{h}(b)] \\ &= [\mathcal{D} \wedge \bar{h}(a)] \wedge \bar{h}(b) - \bar{h}(a) \wedge [\mathcal{D} \wedge \bar{h}(b)] = 0 \end{aligned}$$

Result for position dependent fields is

$$\mathcal{D} \wedge \bar{h}(M) = \bar{h}(\nabla \wedge M)$$

Follows that

$$\mathcal{D} \wedge [\mathcal{D} \wedge \bar{h}(A)] = \mathcal{D} \wedge \bar{h}(\nabla \wedge A) = \bar{h}(\nabla \wedge \nabla \wedge A) = 0$$

Holds for any A . Results in an **algebraic** identity for the Riemann tensor

$$\partial_a \wedge \mathcal{R}(a \wedge b) = 0$$

Trivector $\partial_a \wedge \mathcal{R}(a \wedge b)$ vanishes for all values of **vector** b . 16 constraints, so $\mathcal{R}(B)$ left with 20 degrees of freedom.

SECOND FIELD EQUATION

$\mathcal{R}(B)$ has dimensions $(\text{length})^{-2}$. Ignoring quantum effects, so only have G . Form $\mathcal{R}(B)/G$, has dimensions of **energy density**. Now linear in $\mathcal{R}(B)$, whereas for electromagnetism had quadratic $T(a) = -\frac{1}{2}F a F$.

Expect source for gravitational fields to be matter **stress-energy tensor**. Need a linear function on vectors.

Contract $\mathcal{R}(a \wedge b)$ and define **Ricci tensor**

$$\mathcal{R}(b) = \partial_a \cdot \mathcal{R}(a \wedge b)$$

(NB. Same symbol. Grade of argument distinguishes type.)

Contract again to get **Ricci scalar**

$$\mathcal{R} = \partial_a \cdot \mathcal{R}(a)$$

Our first scalar observable. This is Lagrangian density in formal development.

To get correct equation, use **Jacobi Identity**. One of these for all gauge theories:

$$[D_a, [D_b, D_c]]\psi + [D_c, [D_a, D_b]]\psi + [D_b, [D_c, D_a]]\psi = 0$$

So

$$\mathcal{D}_a \mathcal{R}(b \wedge c) + \mathcal{D}_c \mathcal{R}(a \wedge b) + \mathcal{D}_b \mathcal{R}(c \wedge a) = 0$$

In electromagnetism same as $\nabla \wedge F = 0$. Covariant form in gravity more work. Result is

$$\bar{h}(\partial_a) \wedge [\mathcal{D}_a \mathcal{R}(B) - \mathcal{R}(\mathcal{D}_a B)] = \dot{\mathcal{D}} \wedge \dot{\mathcal{R}}(B) = 0.$$

Called the **Bianchi** identity. Another **useful, practical** result in a **simple, memorable** expression. On contracting Bianchi identity, find that covariantly conserved tensor is the **Einstein tensor**

$$\mathcal{G}(a) = \mathcal{R}(a) - \frac{1}{2} a \mathcal{R}$$

Equate this with the **covariant** matter **stress-energy tensor** $\mathcal{T}(a)$,

$$\mathcal{G}(a) = \kappa \mathcal{T}(a)$$

Constant κ found from spherically symmetric solutions.

Conclusion is $\kappa = 8\pi G$.

TRAJECTORIES AND TANGENTS

Particle follows trajectory $x(\lambda)$ in the STA. Actual path has **no relevance**, only values of fields encountered. What relevance can be attached to the **tangent vector**? After displacement

$$x'(\lambda) = f(x(\lambda))$$

new tangent vector is

$$\partial_\lambda f(x(\lambda)) = \partial_\lambda x(\lambda) \cdot \nabla f(x) = \dot{x}$$

where $\dot{x} = \partial_\lambda x(\lambda)$. Factor of $f(a)$ now. Remove with suitable version of $\bar{h}(a)$. Define

$$v = \mathbf{h}^{-1}(\dot{x}),$$

as the **covariant tangent vector**. Now transforms under rotation gauge

$$v \mapsto v' = Rv\tilde{R}$$

No change to STA trajectory. v can point wherever we like!
Only constraint is v^2 **invariant**. Use this to distinguish spacelike, timelike and null.

POINT PARTICLE EQUATIONS OF MOTION

Proper (invariant) distance along trajectory now

$$s = \int_{\lambda_1}^{\lambda_2} \sqrt{|v^2|} d\lambda.$$

The proper time τ is therefore parameter such that

$$v^2 = 1, \quad \text{where} \quad v = h^{-1}(\dot{x}) = h^{-1}(\partial_\tau x).$$

No fields $\dot{v} = 0$. Make this equation covariant. Have

$\partial_\tau = \dot{x} \cdot \nabla$, so use

$$\partial_\tau v + \Omega(\dot{x}) \cdot v = 0$$

(Can write this more abstractly as $v \cdot \mathcal{D} v = 0$)

$\Omega(\dot{x})$ is a form of **acceleration bivector**. But have to be **careful**.

$\Omega(a)$ is **not gauge invariant**. If transform to $v' = Rv\tilde{R}$ get,

$$\begin{aligned} \partial_\tau v' + \Omega'(\dot{x}) \cdot v' &= \partial_\tau v' + [R\Omega(\dot{x})\tilde{R}] \cdot v' - 2(\dot{R}\tilde{R}) \cdot v' \\ &= Rv \cdot \mathcal{D}v \tilde{R} = 0. \end{aligned}$$

Can add 'acceleration' term $2\dot{R}\tilde{R}$ to $\Omega(a)$. field without altering physics. Can see that

$$\text{gravity} + \text{acceleration} = \text{gravity}'$$

A form of the **equivalence principle**. (Einstein considered lifts accelerating and at rest in a gravitational field.)

Weak Equivalence Principle

the motion of a test particle in a gravitational field is independent of its mass. Automatic — no mass in our equation. Means gravitational mass = inertial masses.

THE METRIC AND GR

Introduce a coordinate frame. Coordinates (x^μ) . Define

$$e_\mu = \partial_\mu x, \quad e^\mu = \nabla x^\mu.$$

Reciprocal to one another. Expand trajectory $x(\lambda)$ in this frame

$$\partial_\lambda x = \partial_\lambda x(x^\mu) = \frac{dx^\mu}{d\lambda} \partial_\mu x = \frac{dx^\mu}{d\lambda} e_\mu.$$

In terms of this the proper distance along a path becomes

$$\begin{aligned} s &= \int_{\lambda_1}^{\lambda_2} \sqrt{|v^2|} d\lambda \\ &= \int_{x_1}^{x_2} |\mathbf{h}^{-1}(e_\mu) \cdot \mathbf{h}^{-1}(e_\nu) dx^\mu dx^\nu|^{1/2} \end{aligned}$$

Comparison with GR $ds^2 = g_{\mu\nu} dx^\mu dx^\nu$ gives metric as

$$g_{\mu\nu} = \mathbf{h}^{-1}(e_\mu) \cdot \mathbf{h}^{-1}(e_\nu).$$

- In GR this is the fundamental object. Gives distance between points on a curved surface. In gauge theory it is derived.

- If $\bar{h}(a)$ and $\Omega(a)$ satisfy gauge field equations then metric solves the GR Einstein equations.
- Minimise proper distance get **geodesic equation**. Same as $v \cdot \mathcal{D}v = 0$.
- Metric independent of rotation gauge. GR never sees this. Gauge nature of GR is **hidden!**
- For metric, displacements look like **coordinate transformations**. Confusing! NB metric cannot be transformed away by change of coordinates.

COVARIANT FRAMES

Define useful frames

$$g_\mu = h^{-1}(e_\mu), \quad g^\mu = \bar{h}(\nabla_{x^\mu}) = \bar{h}(e^\mu).$$

Reciprocal because

$$g_\mu \cdot g^\nu = h^{-1}(e_\mu) \cdot \bar{h}(e^\mu) = e_\mu \cdot e^\nu = \delta_\mu^\nu.$$

Metric now simply

$$g_{\mu\nu} = g_\mu \cdot g_\nu.$$

Also simplify first field equation. Since

$$\mathcal{D} \wedge g^\mu = \mathcal{D} \wedge \bar{h}(e^\mu) = \bar{h}(\nabla \wedge e^\mu)$$

and $\nabla \wedge e^\mu = \nabla \wedge \nabla x^\mu = 0$, can write $\mathcal{S}(a) = 0$ as

$$\mathcal{D} \wedge g^\mu = 0.$$

Now complete the link with GR. We use the abbreviation

$$\mathcal{D}_\mu = e_\mu \cdot \nabla + \Omega(e_\mu) \times = \partial_\mu + \Omega(e_\mu) \times$$

for the covariant derivative in e_μ direction. Act on vector g_μ and express result in $\{g_\mu\}$ frame:

$$\mathcal{D}_\mu g_\nu = \Gamma_{\mu\nu}^\lambda g_\lambda$$

Defines the **Christoffel connection** $\Gamma_{\mu\nu}^\lambda$.

Vectors transform like $\partial_\mu x$ (with $f(a)$) or like ∇ (with $\bar{f}^{-1}(a)$). Mathematicians call these **vectors** and **1-forms**. Like to keep them separate. Identified by the **metric**. We use \bar{h} -field to make all same type — covariant vectors. Then just have rotor group transformations.